Equilibrium in Kinetic Theory

1. Thermalized gas particle in a one-dimensional trap:

(a) Liouville's equation, describing the incompressible nature of phase space density, is

$$\frac{\partial \rho}{\partial t} = -\dot{q}\frac{\partial \rho}{\partial q} - \dot{p}\frac{\partial \rho}{\partial p} = -\frac{\partial \mathcal{H}}{\partial p}\frac{\partial \rho}{\partial q} + \frac{\partial \mathcal{H}}{\partial q}\frac{\partial \rho}{\partial p} \equiv \{\rho, \mathcal{H}\}.$$

For the gas particle confined to a 1-dimensional trap, the Hamiltonian can be written as

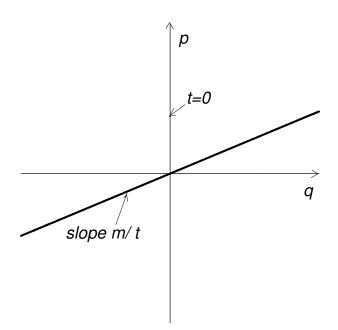
$$\mathcal{H} = \frac{p^2}{2m} + V(q_x) = \frac{p^2}{2m},$$

since $V_{q_x} = 0$, and there is no motion in the y and z directions. With this Hamiltonian, Liouville's equation becomes

$$\frac{\partial \rho}{\partial t} = -\frac{p}{m} \frac{\partial \rho}{\partial q},$$

whose solution, subject to the specified initial conditions, is

$$\rho(q, p, t) = \rho\left(q - \frac{p}{m}t, p, 0\right) = \delta\left(q - \frac{p}{m}t\right)f(p).$$



(b) The expectation value for any observable \mathcal{O} is

$$\langle \mathcal{O} \rangle = \int d\Gamma \mathcal{O} \rho(\Gamma, t),$$

and hence

$$\langle p^2 \rangle = \int p^2 f(p) \delta\left(q - \frac{p}{m}t\right) dp \ dq = \int p^2 f(p) dp$$
$$= \int_{-\infty}^{\infty} dp \ p^2 \frac{1}{\sqrt{2\pi m k_B T}} \exp\left(-\frac{p^2}{2m k_B T}\right) = m k_B T.$$

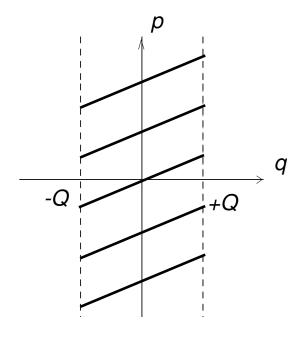
Likewise, we obtain

$$\left\langle q^2 \right\rangle = \int q^2 f(p) \delta\left(q - \frac{p}{m}t\right) dp \, dq = \int \left(\frac{p}{m}t\right)^2 f(p) dp = \left(\frac{t}{m}\right)^2 \int p^2 f(p) dp = \frac{k_B T}{m}t^2.$$

(c) Now suppose that hard walls are placed at $q = \pm Q$. The appropriate relaxation time τ , is the characteristic length between the containing walls divided by the characteristic velocity of the particle, i.e.

$$\tau \sim \frac{2Q}{|\dot{q}|} = \frac{2Qm}{\sqrt{\langle p^2 \rangle}} = 2Q\sqrt{\frac{m}{k_BT}}.$$

Initially $\rho(q, p, t)$ resembles the distribution shown in part (a), but each time the particle hits the barrier, reflection changes p to -p. As time goes on, the slopes become less, and $\rho(q, p, t)$ becomes a set of closely spaced lines whose separation vanishes as 2mQ/t.



(d) We can choose any resolution ε in (p,q) space, subdividing the plane into an array of pixels of this area. For any ε , after sufficiently long time many lines will pass through this area. Averaging over them leads to

$$\tilde{\rho}(q, p, t \gg \tau) = \frac{1}{2Q} f(p),$$

as (i) the density f(p) at each p is always the same, and (ii) all points along $q \in [-Q, +Q]$ are equally likely. For the time variation of this coarse-grained density, we find

$$\frac{\partial \tilde{\rho}}{\partial t} = -\frac{p}{m} \frac{\partial \tilde{\rho}}{\partial q} = 0, \quad \text{i.e.} \quad \tilde{\rho} \quad \text{is stationary.}$$

2. Entropy and the Ensemble Density:

(a) A candidate "entropy" is defined by

$$S(t) = -\int d\Gamma \rho(\Gamma, t) \ln \rho(\Gamma, t) = -\langle \ln \rho(\Gamma, t) \rangle.$$

Taking the derivative with respect to time gives

$$\frac{dS}{dt} = -\int d\Gamma \left(\frac{\partial\rho}{\partial t} \ln\rho + \rho \frac{1}{\rho} \frac{\partial\rho}{\partial t}\right) = -\int d\Gamma \frac{\partial\rho}{\partial t} \left(\ln\rho + 1\right).$$

Substituting the expression for $\partial \rho / \partial t$ obtained from Liouville's theorem gives

$$\frac{dS}{dt} = -\int d\Gamma \sum_{i=1}^{3N} \left(\frac{\partial \rho}{\partial p_i} \frac{\partial \mathcal{H}}{\partial q_i} - \frac{\partial \rho}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i} \right) \left(\ln \rho + 1 \right).$$

(Here the index *i* is used to label the 3 coordinates, as well as the *N* particles, and hence runs from 1 to 3N.) Integrating the above expression by parts yields[†]

[†] This is standard integration by parts, i.e. $\int_{b}^{a} F dG = FG|_{b}^{a} - \int_{b}^{a} G dF$. Looking explicitly at one term in the expression to be integrated in this problem,

$$\int \prod_{i=1}^{3N} dV_i \frac{\partial \rho}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i} = \int dq_1 dp_1 \cdots dq_i dp_i \cdots dq_{3N} dp_{3N} \frac{\partial \rho}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i},$$

we identify $dG = dq_i \frac{\partial \rho}{\partial q_i}$, and F with the remainder of the expression. Noting that $\rho(q_i) = 0$ at the boundaries of the box, we get

$$\int \prod_{i=1}^{3N} dV_i \frac{\partial \rho}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i} = -\int \prod_{i=1}^{3N} dV_i \rho \frac{\partial}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i}$$

$$\begin{split} \frac{dS}{dt} &= \int d\Gamma \sum_{i=1}^{3N} \left[\rho \frac{\partial}{\partial p_i} \left(\frac{\partial \mathcal{H}}{\partial q_i} \left(\ln \rho + 1 \right) \right) - \rho \frac{\partial}{\partial q_i} \left(\frac{\partial \mathcal{H}}{\partial p_i} \left(\ln \rho + 1 \right) \right) \right] \\ &= \int d\Gamma \sum_{i=1}^{3N} \left[\rho \frac{\partial^2 \mathcal{H}}{\partial p_i \partial q_i} (\ln \rho + 1) + \rho \frac{\partial \mathcal{H}}{\partial q_i} \frac{1}{\rho} \frac{\partial \rho}{\partial p_i} - \rho \frac{\partial^2 \mathcal{H}}{\partial q_i \partial p_i} (\ln \rho + 1) - \rho \frac{\partial \mathcal{H}}{\partial p_i} \frac{1}{\rho} \frac{\partial \rho}{\partial q_i} \right] \\ &= \int d\Gamma \sum_{i=1}^{3N} \left[\frac{\partial \mathcal{H}}{\partial q_i} \frac{\partial \rho}{\partial p_i} - \frac{\partial \mathcal{H}}{\partial p_i} \frac{\partial \rho}{\partial q_i} \right]. \end{split}$$

Integrating the final expression by parts gives

$$\frac{dS}{dt} = -\int d\Gamma \sum_{i=1}^{3N} \left[-\rho \frac{\partial^2 \mathcal{H}}{\partial p_i \partial q_i} + \rho \frac{\partial^2 \mathcal{H}}{\partial q_i \partial p_i} \right] = 0.$$

(b) There are two constraints, normalization and constant average energy, written respectively as

$$\int d\Gamma \rho(\Gamma) = 1$$
, and $\langle \mathcal{H} \rangle = \int d\Gamma \rho(\Gamma) \mathcal{H} = E$.

Rewriting the expression for entropy,

$$S(t) = \int d\Gamma \rho(\Gamma) \left[-\ln \rho(\Gamma) - \alpha - \beta \mathcal{H} \right] + \alpha + \beta E,$$

where α and β are Lagrange multipliers used to enforce the two constraints. Extremizing the above expression with respect to the function $\rho(\Gamma)$, results in

$$\frac{\partial S}{\partial \rho(\Gamma)}\Big|_{\rho=\rho_{max}} = -\ln \rho_{max}(\Gamma) - \alpha - \beta \mathcal{H}(\Gamma) - 1 = 0.$$

The solution to this equation is

$$\ln \rho_{max} = -(\alpha + 1) - \beta \mathcal{H},$$

which can be rewritten as

$$\rho_{max} = C \exp(-\beta \mathcal{H}), \quad \text{where} \quad C = e^{-(\alpha+1)}.$$

(c) The density obtained in part (b) is stationary, as can be easily checked from

$$\frac{\partial \rho_{max}}{\partial t} = \{\rho_{max}, \mathcal{H}\} = \left\{Ce^{-\beta\mathcal{H}}, \mathcal{H}\right\}$$
$$= \frac{\partial \mathcal{H}}{\partial p}C(-\beta)\frac{\partial \mathcal{H}}{\partial q}e^{-\beta\mathcal{H}} - \frac{\partial \mathcal{H}}{\partial q}C(-\beta)\frac{\partial \mathcal{H}}{\partial p}e^{-\beta\mathcal{H}} = 0.$$

(d) Liouville's equation preserves the information content of the PDF $\rho(\Gamma, t)$, and hence S(t) does not increase in time. However, as illustrated in the example in problem 1, the density becomes more finely dispersed in phase space. In the presence of any coarsegraining of phase space, information disappears. The maximum entropy, corresponding to $\tilde{\rho}$, describes equilibrium in this sense.

3. The Vlasov equation: The BBGKY hierarchy

$$\begin{bmatrix} \frac{\partial}{\partial t} + \sum_{n=1}^{s} \frac{\vec{p}_{n}}{m} \cdot \frac{\partial}{\partial \vec{q}_{n}} - \sum_{n=1}^{s} \left(\frac{\partial U}{\partial \vec{q}_{n}} + \sum_{l} \frac{\partial \mathcal{V}(\vec{q}_{n} - \vec{q}_{l})}{\partial \vec{q}_{n}} \right) \cdot \frac{\partial}{\partial \vec{p}_{n}} \end{bmatrix} f_{s}$$
$$= \sum_{n=1}^{s} \int dV_{s+1} \frac{\partial \mathcal{V}(\vec{q}_{n} - \vec{q}_{s+1})}{\partial \vec{q}_{n}} \cdot \frac{\partial f_{s+1}}{\partial \vec{p}_{n}},$$

has the characteristic time scales

$$\begin{cases} \frac{1}{\tau_U} \sim \frac{\partial U}{\partial \vec{q}} \cdot \frac{\partial}{\partial \vec{p}} \sim \frac{v}{L}, \\ \frac{1}{\tau_c} \sim \frac{\partial \mathcal{V}}{\partial \vec{q}} \cdot \frac{\partial}{\partial \vec{p}} \sim \frac{v}{\lambda}, \\ \frac{1}{\tau_X} \sim \int dx \frac{\partial \mathcal{V}}{\partial \vec{q}} \cdot \frac{\partial}{\partial \vec{p}} \frac{f_{s+1}}{f_s} \sim \frac{1}{\tau_c} \cdot n\lambda^3, \end{cases}$$

where $n\lambda^3$ is the number of particles within the interaction range λ , and v is a typical velocity. The Boltzmann equation is obtained in the dilute limit, $n\lambda^3 \ll 1$, by disregarding terms of order $1/\tau_X \ll 1/\tau_c$. The Vlasov equation is obtained in the dense limit of $n\lambda^3 \gg 1$ by ignoring terms of order $1/\tau_c \ll 1/\tau_X$).

(a) Let bfx_i denote the coordinates and momenta for particle *i*. Starting from the joint probability $\rho_N = \prod_{i=1}^N \rho_1(\mathbf{x}_i, t)$, for *independent particles*, we find

$$f_s = \frac{N!}{(N-s)!} \int \prod_{\alpha=s+1}^N dV_\alpha \rho_N = \frac{N!}{(N-s)!} \prod_{n=1}^s \rho_1(\mathbf{x}_n, t).$$

The normalizations follow from

$$\int d\Gamma \rho = 1, \quad \Longrightarrow \quad \int dV_1 \rho_1(\mathbf{x}, t) = 1,$$

and

$$\int \prod_{n=1}^{s} dV_n f_s = \frac{N!}{(N-s)!} \approx N^s \quad \text{for} \quad s \ll N.$$

(b) Noting that

$$\frac{f_{s+1}}{f_s} = \frac{(N-s)!}{(N-s-1)!} \rho_1(\mathbf{x}_{s+1}),$$

the reduced BBGKY hierarchy is

$$\left[\frac{\partial}{\partial t} + \sum_{n=1}^{s} \left(\frac{\vec{p}_{n}}{m} \cdot \frac{\partial}{\partial \vec{q}_{n}} - \frac{\partial U}{\partial \vec{q}_{n}} \cdot \frac{\partial}{\partial \vec{p}_{n}}\right)\right] f_{s}$$

$$\approx \sum_{n=1}^{s} \int dV_{s+1} \frac{\partial \mathcal{V}(\vec{q}_{n} - \vec{q}_{s+1})}{\partial \vec{q}_{n}} \cdot \frac{\partial}{\partial \vec{p}_{n}} \left[(N-s)f_{s}\rho_{1}(\mathbf{x}_{s+1})\right]$$

$$\approx \sum_{n=1}^{s} \frac{\partial}{\partial \vec{q}_{n}} \left[\int dV_{s+1}\rho_{1}(\mathbf{x}_{s+1})\mathcal{V}(\vec{q}_{n} - \vec{q}_{s+1}) \cdot N\right] \frac{\partial}{\partial \vec{p}_{n}} f_{s},$$

where we have used the approximation $(N - s) \approx N$ for $N \gg s$. Rewriting the above expression,

$$\left[\frac{\partial}{\partial t} + \sum_{n=1}^{s} \left(\frac{\vec{p}_n}{m} \cdot \frac{\partial}{\partial \vec{q}_n} - \frac{\partial U_{eff}}{\partial \vec{q}_n} \cdot \frac{\partial}{\partial \vec{p}_n}\right)\right] f_s = 0,$$

where

$$U_{eff} = U(\vec{q}) + N \int dV' \mathcal{V}(\vec{q} - \vec{q}') \rho_1(\mathbf{x}', t).$$

(c) Starting from

$$\rho_1 = g(\vec{p})/V,$$

we obtain

$$\mathcal{H}_{eff} = \sum_{i=1}^{N} \left[\frac{\vec{p_i}^2}{2m} + U_{eff}(\vec{q_i}) \right],$$

with

$$U_{eff} = 0 + N \int dV' \mathcal{V}(\vec{q} - \vec{q}') \frac{1}{V} g(\vec{p}) = \frac{N}{V} \int d^3q \mathcal{V}(\vec{q}).$$

(We have taken advantage of the normalization $\int d^3pg(\vec{p}) = 1$.) Substituting into the Vlasov equation yields

$$\left(\frac{\partial}{\partial t} + \frac{\vec{p}}{m} \cdot \frac{\partial}{\partial \vec{q}}\right)\rho_1 = 0.$$

There is no relaxation towards equilibrium because there are no collisions which allow $g(\vec{p})$ to relax. The momentum of each particle is conserved by \mathcal{H}_{eff} ; i.e. $\{\rho_1, \mathcal{H}_{eff}\} = 0$, preventing its change.

4. Two component plasma:

(a) The Hamiltonian for the two component mixture is

$$\mathcal{H} = \sum_{i=1}^{N} \left[\frac{\vec{p}_i^2}{2m_+} + \frac{\vec{p}_i^2}{2m_-} \right] + \sum_{i,j=1}^{2N} e_i e_j \frac{1}{|\vec{q}_i - \vec{q}_j|} + \sum_{i=1}^{2N} e_i \Phi_{ext}(\vec{q}_i),$$

where $C(\vec{q}_i - \vec{q}_j) = 1/|\vec{q}_i - \vec{q}_j|$, resulting in

$$\frac{\partial \mathcal{H}}{\partial \vec{q_i}} = e_i \frac{\partial \Phi_{ext}}{\partial \vec{q_i}} + e_i \sum_{j \neq i} e_j \frac{\partial}{\partial \vec{q_i}} C(\vec{q_i} - \vec{q_j}).$$

Substituting this into the Vlasov equation, we obtain

$$\begin{cases} \left[\frac{\partial}{\partial t} + \frac{\vec{p}}{m_{+}} \cdot \frac{\partial}{\partial \vec{q}} + e\frac{\partial \Phi_{eff}}{\partial \vec{q}} \cdot \frac{\partial}{\partial \vec{p}}\right] f_{+}(\vec{p},\vec{q},t) = 0, \\ \left[\frac{\partial}{\partial t} + \frac{\vec{p}}{m_{-}} \cdot \frac{\partial}{\partial \vec{q}} - e\frac{\partial \Phi_{eff}}{\partial \vec{q}} \cdot \frac{\partial}{\partial \vec{p}}\right] f_{-}(\vec{p},\vec{q},t) = 0. \end{cases}$$

(b) Setting $f_{\pm}(\vec{p}, \vec{q}) = g_{\pm}(\vec{p})n_{\pm}(\vec{q})$, and using $\int d^3p g_{\pm}(\vec{p}) = 1$, the integrals in the effective potential simplify to

$$\Phi_{eff}(\vec{q},t) = \Phi_{ext}(\vec{q}\,) + e \int d^3 q' C(\vec{q}-\vec{q}\,') \left[n_+(\vec{q}\,') - n_-(\vec{q}\,')\right].$$

Apply ∇^2 to the above equation, and use $\nabla^2 \Phi_{ext} = 4\pi \rho_{ext}$ and $\nabla^2 C(\vec{q} - \vec{q'}) = 4\pi \delta^3(\vec{q} - \vec{q'})$, to obtain

$$\nabla^2 \phi_{eff} = 4\pi \rho_{ext} + 4\pi e \left[n_+(\vec{q}) - n_-(\vec{q}) \right].$$

(c) Linearizing the Boltzmann weights gives

$$n_{\pm} = n_o \exp[\mp \beta e \Phi_{eff}(\vec{q}\,)] \approx n_o \left[1 \mp \beta e \Phi_{eff}\right],$$

resulting in

$$\nabla^2 \Phi_{eff} = 4\pi \rho_{ext} + \frac{1}{\lambda^2} \Phi_{eff},$$

with the screening length given by

$$\lambda^2 = \frac{k_B T}{8\pi n_o e^2}.$$

(d) We want to show that the Debye equation has the general solution

$$\Phi_{eff}(\vec{q}\,) = \int d^3 \vec{q} G(\vec{q} - \vec{q}\,') \rho_{ext}(\vec{q}\,'),$$

where

$$G(\vec{q}) = \frac{\exp(-|q|/\lambda)}{|q|}.$$

Effectively, we want to show that $\nabla^2 G = G/\lambda^2$ for $\vec{q} \neq 0$. In spherical coordinates, $G = \exp(-r/\lambda)/r$. Evaluating ∇^2 in spherical coordinates gives

$$\nabla^2 G = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial G}{\partial r} \right) = \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \left[-\frac{1}{\lambda} \frac{e^{-r/\lambda}}{r} - \frac{e^{-r/\lambda}}{r^2} \right]$$
$$= \frac{1}{r^2} \left[\frac{1}{\lambda^2} r e^{-r/\lambda} - \frac{1}{\lambda} e^{-r/\lambda} + \frac{1}{\lambda} e^{-r/\lambda} \right] = \frac{1}{\lambda^2} \frac{e^{-r/\lambda}}{r} = \frac{G}{\lambda^2}.$$

(e) The Vlasov equation assumes the limit $n_o \lambda^3 \gg 1$, which requires that

$$\frac{(k_B T)^{3/2}}{n_o^{1/2} e^3} \gg 1, \quad \Longrightarrow \quad \frac{e^2}{k_B T} \ll n_o^{-1/3} \sim \ell,$$

where ℓ is the interparticle spacing. In terms of the interparticle spacing, the selfconsistency condition is

$$\frac{e^2}{\ell} \ll k_B T,$$

i.e. the interaction energy is much less than the kinetic (thermal) energy.

(f) A characteristic time is obtained from

$$\tau \sim \frac{\lambda}{c} \sim \sqrt{\frac{k_B T}{n_o e^2}} \cdot \sqrt{\frac{m}{k_B T}} \sim \sqrt{\frac{m}{n_o e^2}} \sim \frac{1}{\omega_p},$$

where ω_p is the plasma frequency.
