

III.G Conservation Laws

• *Approach to equilibrium:* We now address the third question posed in the introduction of how the gas reaches its final equilibrium. Consider a situation in which the gas is perturbed from the equilibrium form described by eq.(III.53), and follow its relaxation to equilibrium. There is a hierarchy of mechanisms that operate at different time scales.

(i) The fastest processes are the two body collisions of particles in immediate vicinity.

Over a time scale of the order of τ_c , $f_2(\vec{q}_1, \vec{q}_2, t)$ relaxes to $f_1(\vec{q}_1, t)f_1(\vec{q}_2, t)$ for separations $|\vec{q}_1 - \vec{q}_2| \gg d$. Similar relaxations occur for the higher order densities f_s .

(ii) At the next stage, f_1 relaxes to a *local equilibrium* form, as in eq.(III.50), over the time scale of the mean free time τ_\times . This is the intrinsic scale set by the collision term on the right hand side of the Boltzmann equation. After this time interval, at each point we can define a local (time dependent) density by integrating over all momenta as

$$n(\vec{q}, t) = \int d^3\vec{p} f_1(\vec{p}, \vec{q}, t), \quad (\text{III.66})$$

as well as a local expectation value for any operator $\mathcal{O}(\vec{p}, \vec{q}, t)$

$$\langle \mathcal{O}(\vec{q}, t) \rangle = \frac{1}{n(\vec{q}, t)} \int d^3\vec{p} f_1(\vec{p}, \vec{q}, t) \mathcal{O}(\vec{p}, \vec{q}, t). \quad (\text{III.67})$$

(iii) After the densities and expectation values have relaxed to their local equilibrium forms in the intrinsic time scales τ_c and τ_\times , there is a subsequent relaxation to equilibrium over extrinsic time and length scales. The slow relaxation is controlled by the *conserved quantities*, which evolve according to *hydrodynamic equations*.

Conserved quantities, are left unchanged by the two body collisions, i.e. satisfy

$$\chi(\vec{p}_1, \vec{q}, t) + \chi(\vec{p}_2, \vec{q}, t) = \chi(\vec{p}_1', \vec{q}, t) + \chi(\vec{p}_2', \vec{q}, t), \quad (\text{III.68})$$

where (\vec{p}_1, \vec{p}_2) and (\vec{p}_1', \vec{p}_2') refer to the momenta before and after a collision respectively. For such quantities, we have

$$J = \int d^3\vec{p} \chi(\vec{p}, \vec{q}, t) \left. \frac{df_1}{dt} \right|_{\text{coll.}} = 0. \quad (\text{III.69})$$

• **Proof:** Using the form of the collision integral, we have

$$J = \int d^3\vec{p}_1 d^3\vec{p}_2 d^2\vec{b} |\vec{v}_1 - \vec{v}_2| [f_1(\vec{p}_1)f_1(\vec{p}_2) - f_1(\vec{p}_1')f_1(\vec{p}_2')] \chi(\vec{p}_1). \quad (\text{III.70})$$

We now perform the same set of changes of variables that were used in the proof of the H-theorem. The first step is averaging after exchange of the dummy variables \vec{p}_1 and \vec{p}_2 , leading to

$$J = \frac{1}{2} \int d^3\vec{p}_1 d^3\vec{p}_2 d^2\vec{b} |\vec{v}_1 - \vec{v}_2| [f_1(\vec{p}_1)f_1(\vec{p}_2) - f_1(\vec{p}_1')f_1(\vec{p}_2')] (\chi(\vec{p}_1) + \chi(\vec{p}_2)). \quad (\text{III.71})$$

Next, change variables from the originators $(\vec{p}_1, \vec{p}_2, \vec{b})$, to the products $(\vec{p}_1', \vec{p}_2', \vec{b}')$ of the collision. After relabeling the integration variables, the above equation is transformed to

$$J = \frac{1}{2} \int d^3\vec{p}_1 d^3\vec{p}_2 d^2\vec{b} |\vec{v}_1 - \vec{v}_2| [f_1(\vec{p}_1')f_1(\vec{p}_2') - f_1(\vec{p}_1)f_1(\vec{p}_2)] (\chi(\vec{p}_1') + \chi(\vec{p}_2')). \quad (\text{III.72})$$

Averaging the last two equations leads to

$$J = \frac{1}{4} \int d^3\vec{p}_1 d^3\vec{p}_2 d^2\vec{b} |\vec{v}_1 - \vec{v}_2| [f_1(\vec{p}_1)f_1(\vec{p}_2) - f_1(\vec{p}_1')f_1(\vec{p}_2')] [\chi(\vec{p}_1) + \chi(\vec{p}_2) - \chi(\vec{p}_1') - \chi(\vec{p}_2')], \quad (\text{III.73})$$

which is zero from eq.(III.68).

Let us explore the consequences of this result for the evolution of expectation values involving χ . Substituting for the collision term in eq.(III.69) the streaming terms on the left hand side of the Boltzmann equation leads to

$$J = \int d^3\vec{p} \chi(\vec{p}, \vec{q}, t) \left[\partial_t + \frac{p_\alpha}{m} \partial_\alpha + F_\alpha \frac{\partial}{\partial p_\alpha} \right] f_1 = 0, \quad (\text{III.74})$$

where we have introduced the notations $\partial_t \equiv \partial/\partial t$, $\partial_\alpha \equiv \partial/\partial q_\alpha$, and $F_\alpha = -\partial U/\partial q_\alpha$. We can manipulate the above equation into the form

$$\int d^3\vec{p} \left\{ \left[\partial_t + \frac{p_\alpha}{m} \partial_\alpha + F_\alpha \frac{\partial}{\partial p_\alpha} \right] (\chi f_1) - f \left[\partial_t + \frac{p_\alpha}{m} \partial_\alpha + F_\alpha \frac{\partial}{\partial p_\alpha} \right] \chi \right\} = 0. \quad (\text{III.75})$$

The third term is zero, as it is a complete derivative. Using the definition of expectation values in eq.(III.67), the remaining terms can be rearranged into

$$\partial_t (n \langle \chi \rangle) + \partial_\alpha \left(n \left\langle \frac{p_\alpha}{m} \chi \right\rangle \right) - n \langle \partial_t \chi \rangle - n \left\langle \frac{p_\alpha}{m} \partial_\alpha \chi \right\rangle - n F_\alpha \left\langle \frac{\partial \chi}{\partial p_\alpha} \right\rangle = 0. \quad (\text{III.76})$$

As discussed earlier, for elastic collisions, there are 5 conserved quantities: particle number, the three components of momentum, and kinetic energy. Each leads to a corresponding hydrodynamic equation, as constructed below:

(a) *Particle number*: Setting $\chi = 1$ in eq.(III.76) leads to

$$\partial_t n + \partial_\alpha (n u_\alpha) = 0, \quad (\text{III.77})$$

where we have introduced the local velocity

$$\vec{u} \equiv \left\langle \frac{\vec{p}}{m} \right\rangle. \quad (\text{III.78})$$

This equation simply states that the time variation of the local particle density is due to a particle current $\vec{J}_n = n\vec{u}$.

(b) *Momentum*: Any linear function of the momentum \vec{p} is conserved in the collision, and we shall explore the consequences of the conservation of

$$\vec{c} \equiv \frac{\vec{p}}{m} - \vec{u}. \quad (\text{III.79})$$

Substituting c_α into eq.(III.76) leads to

$$\partial_\beta (n \langle (u_\beta + c_\beta) c_\alpha \rangle) + n \partial_t u_\alpha + n \partial_\beta u_\alpha \langle u_\beta + c_\beta \rangle - n \frac{F_\alpha}{m} = 0. \quad (\text{III.80})$$

Taking advantage of $\langle c_\alpha \rangle = 0$, from eqs.(III.78) and (III.79), leads to

$$\partial_t u_\alpha + u_\beta \partial_\beta u_\alpha = \frac{F_\alpha}{m} - \frac{1}{mn} \partial_\beta P_{\alpha\beta}, \quad (\text{III.81})$$

where we have introduced the *pressure tensor*,

$$P_{\alpha\beta} \equiv mn \langle c_\alpha c_\beta \rangle. \quad (\text{III.82})$$

The left hand side of the equation is the acceleration of an element of the fluid $d\vec{u}/dt$, which should equal \vec{F}_{net}/m according to Newton's equation. Clearly the net force has acquired an additional component due to the variations in the pressure tensor in the fluid.

(c) *Kinetic energy*: We first introduce an average local kinetic energy

$$\varepsilon \equiv \left\langle \frac{mc^2}{2} \right\rangle = \left\langle \frac{p^2}{2m} - \vec{p} \cdot \vec{u} + \frac{mu^2}{2} \right\rangle, \quad (\text{III.83})$$

and then examine the conservation law obtained by setting χ equal to $mc^2/2$ in eq.(III.76). Noting that $\partial\chi = mc_\beta \partial c_\beta$, we obtain

$$\partial_t (n\varepsilon) + \partial_\alpha \left(n \left\langle (u_\alpha + c_\alpha) \frac{mc^2}{2} \right\rangle \right) + nm \partial_t u_\beta \langle c_\beta \rangle + nm \partial_\alpha u_\beta \langle (u_\alpha + c_\alpha) c_\beta \rangle - n F_\alpha \langle c_\alpha \rangle = 0. \quad (\text{III.84})$$

Taking advantage of $\langle c_\alpha \rangle = 0$, the above equation is simplified to

$$\partial_t(n\varepsilon) + \partial_\alpha(nu_\alpha\varepsilon) + \partial_\alpha\left(n\left\langle c_\alpha\frac{mc^2}{2}\right\rangle\right) + P_{\alpha\beta}\partial_\alpha u_\beta = 0. \quad (\text{III.85})$$

We next take out the dependence on n in the first two terms of the above equation, finding

$$\varepsilon\partial_t n + n\partial_t\varepsilon + \varepsilon\partial_\alpha(nu_\alpha) + nu_\alpha\partial_\alpha\varepsilon + \partial_\alpha h_\alpha + P_{\alpha\beta}u_{\alpha\beta} = 0, \quad (\text{III.86})$$

where we have also introduced the local *heat flux*

$$\vec{h} \equiv \frac{nm}{2} \langle c_\alpha c^2 \rangle, \quad (\text{III.87})$$

and the *rate of strain tensor*

$$u_{\alpha\beta} = \frac{1}{2}(\partial_\alpha u_\beta + \partial_\beta u_\alpha). \quad (\text{III.88})$$

Eliminating the first and third terms in eq.(III.86) with the aid of eq.(III.77) leads to

$$\partial_t\varepsilon + u_\alpha\partial_\alpha\varepsilon = -\frac{1}{n}\partial_\alpha h_\alpha - \frac{1}{n}P_{\alpha\beta}u_{\alpha\beta}. \quad (\text{III.89})$$

Clearly to solve the hydrodynamic equations for n , \vec{u} , and ε , we need expressions for $P_{\alpha\beta}$ and \vec{h} , which are either given phenomenologically, or calculated from the density f_1 , as in the next sections.